
Modern approaches to quantum gravity

Solution 8

Fall 2025

1. Unitarity bounds

(a) Let us illustrate the solution with the scalar. As you've seen in the lecture.

$$\langle O|K_\mu P_\nu|O\rangle = 2\Delta\delta_{\mu\nu}\langle O|O\rangle = 2\Delta\delta_{\mu\nu}. \quad (1)$$

Thus, the norm of the state $|z\rangle = z_\mu P^\mu|O\rangle$ is $\langle z|z\rangle = 2\Delta|z|^2$. Since this must be positive, we conclude that $\Delta \geq 0$, with $\Delta = 0$ only possible when $P_\mu|O\rangle = 0$. This represents a state without descendants, meaning that $[P_\mu, O(x)] = i\partial_\mu O(x) = 0$, so $O(x)$ does not depend on its position x^μ . Thus, $O(x)$ must be a constant operator. In fact, this implies that $O(x)$ is the identity operator, corresponding to the vacuum state (which transforms as a scalar primary with $\Delta = 0$).

(b) For tensor fields (point (a) and (c) of the exercise set), let us define

$$|\mathcal{O}_\mu[k]\rangle := |\mathcal{O}_{\mu\sigma_1\dots\sigma_{\ell-1}}\rangle k^{\sigma_1\dots\sigma_{\ell-1}}, \quad |k\rangle := P_\mu|\mathcal{O}^\mu[k]\rangle, \quad (2)$$

where $k^{\sigma_1\dots\sigma_{\ell-1}}$ is an auxiliary tensor that can be assumed to be completely symmetric and traceless. We seek the consequences of $\langle k|k\rangle \geq 0$. Hence, we compute using the conformal algebra

$$0 \leq \langle k|k\rangle = \langle \mathcal{O}_\mu[k]|K^\mu P^\nu|\mathcal{O}_\nu[k]\rangle = 2\Delta\langle \mathcal{O}_\mu[k]|\mathcal{O}^\mu[k]\rangle - 2i\langle \mathcal{O}_\mu[k]|M^{\mu\nu}|\mathcal{O}_\nu[k]\rangle. \quad (3)$$

Next, we compute using the explicit definition of $\hat{L}_{\mu\nu}$ and $(M_{\mu\nu})^\alpha_\beta$ that

$$\hat{L}^{\mu\nu}|\mathcal{O}_\nu[k]\rangle = -i(\ell + d - 2)|\mathcal{O}_\mu[k]\rangle. \quad (4)$$

Hence,

$$0 \leq \langle k|k\rangle = 2[\Delta - \ell - d + 2]\langle \mathcal{O}_\mu[k]|\mathcal{O}^\mu[k]\rangle. \quad (5)$$

Since $\langle \mathcal{O}_\mu[k]|\mathcal{O}^\mu[k]\rangle \geq 0$, we conclude that $\Delta \geq \ell + d - 2$, which is the desired bound.

If the norm of $|k\rangle$ vanishes, then $|k\rangle$ itself vanishes, implying $P^\mu|\mathcal{O}_{\mu\sigma_1\dots\sigma_{\ell-1}}\rangle = 0$. Repeating the previous argument, we conclude that

$$\partial_\mu \mathcal{O}^{\mu\sigma_1\dots\sigma_{\ell-1}}(x) = 0, \quad (6)$$

meaning that \mathcal{O} is conserved. Indeed, we already saw that a Noether current has dimension $d - 1$ and the stress tensor has dimension d ; here we have proven the converse.

- (c) Checking that \hat{C} commutes with all generators can be done by hand or in Mathematica straightforwardly. To evaluate its action on a primary state $|A_{\mu_1 \dots \mu_\ell}\rangle$ of dimension Δ and spin ℓ , we note that

$$K_\mu P^\mu = P_\mu K^\mu + 2dD, \quad (7)$$

since $\delta^{\mu\nu} M_{\mu\nu} = 0$ (as $M_{\mu\nu}$ is antisymmetric). Thus,

$$\hat{C}|\text{primary}\rangle = [D^2 - dD + C_{\text{SO}(d)}] |\text{primary}\rangle. \quad (8)$$

In particular, if the primary has spin ℓ and dimension Δ , we find that

$$\hat{C}|A_{\mu_1 \dots \mu_\ell}\rangle = [\Delta(\Delta - d) + \ell(\ell + d - 2)] |A_{\mu_1 \dots \mu_\ell}\rangle. \quad (9)$$

2. Radial Quantization of the Free Scalar

- (a) Let's introduce some notation. Let

$$ds_{S^{d-1}}^2 = h_{ij}(\theta^k) d\theta^i d\theta^j$$

be the metric of the unit sphere S^{d-1} . For instance, in $d = 2, 3$ we have

$$ds_{S^1}^2 = d\phi^2, \quad ds_{S^2}^2 = d\theta^2 + \sin^2(\theta) d\phi^2.$$

The corresponding volume element is $\sqrt{h} \prod_i d\theta^i \equiv d\Omega$, to connect with the notation from the exercise. The metric in radial coordinates (r, θ^i) resp. cylinder coordinates (σ, θ^i) coordinates reads

$$ds^2 = dr^2 + r^2 ds_{S^{d-1}}^2 = e^{2\sigma} [d\sigma^2 + ds_{S^{d-1}}^2]$$

since $dr = d(e^\sigma) = e^\sigma d\sigma$. In particular, in the (σ, θ^i) coordinates we have $\sqrt{g} = \exp(d\sigma)\sqrt{h}$ so the volume element is $e^{d\sigma} d\sigma d\Omega$.

After throwing away a boundary term, the action is given by

$$S[\phi] = \int d^d x \frac{1}{2} (-\phi \square \phi), \quad \square = \partial_\mu \partial^\mu$$

working in Cartesian coordinates x^μ . In a general coordinate system, the Laplacian acting on a general function f is instead

$$\square f = \frac{1}{\sqrt{g}} \partial_\mu (\sqrt{g} g^{\mu\nu} \partial_\nu f)$$

Evaluating the Laplacian in cylinder coordinates (σ, θ^i) and plugging it into the above relation will yield the correct answer. Nevertheless, we will use a different strategy. The Euclidean action can be rewritten in a purely covariant form as

$$S_E = \int d^d x \frac{1}{2} (\partial_\mu \phi)^2 = \frac{1}{2} \int d^d x \sqrt{g} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi$$

We can now just use our change of variables as

$$\begin{aligned} S_E &= \frac{1}{2} \int d\sigma \int d^{d-1}\theta \sqrt{h} e^{(d-2)\sigma} [(\partial_\sigma \phi)^2 + h^{ij} \partial_i \phi \partial_j \phi] \\ &= \frac{1}{2} \int d\sigma d\Omega e^{(d-2)\sigma} [(\partial_\sigma \phi)^2 + h^{ij} \partial_i \phi \partial_j \phi] \end{aligned}$$

According to the exercise, we should define a field $\phi = \exp\left(-\frac{1}{2}(d-2)\sigma\right) \chi$. With this definition, we have

$$\partial_\sigma \phi = \frac{2-d}{2} e^{-\frac{d-2}{2}\sigma} \chi + e^{-\frac{d-2}{2}\sigma} \partial_\sigma \chi$$

Plugging that into the action and neglecting total derivatives yields the answer from the exercise.

Using the other strategy, we obtain

$$-e^{d\sigma} \phi \square \phi = \chi D\chi, \quad D = -\frac{\partial^2}{\partial \sigma^2} + \left(\frac{d-2}{2}\right)^2 - \square_{S^{d-1}}.$$

Therefore

$$S[\phi] = \frac{1}{2} \int d\sigma d\Omega \left[(\partial_\sigma \chi)^2 - \chi \square_{S^{d-1}} \chi + \left(\frac{d-2}{2}\right)^2 \chi^2 \right].$$

If we write $-\chi \square_{S^{d-1}} \chi = \partial_i \chi \partial^i \chi$, we recover the result from the exercise.

(b) **Solutions of the Equation $D\chi = 0$**

Suppose that $Y(\theta^i)$ is an eigenfunction of the Laplacian $\square_{S^{d-1}}$ in d dimensions, with eigenvalue $-\gamma$. If we take the Ansatz

$$\chi(\sigma, \theta^i) = e^{\pm \omega \sigma} Y(\theta^i),$$

then

$$D\chi = \left[-\omega^2 + \left(\frac{d-2}{2}\right)^2 + \gamma \right] \chi \implies \omega = \sqrt{\gamma + \left(\frac{d-2}{2}\right)^2}.$$

If you have studied the theory of spherical harmonics, you may know that in d dimensions, the spherical harmonics of S^{d-1} have eigenvalues $\gamma = \ell(\ell + d - 2)$, so the possible values of ω are

$$\omega_\ell = \sqrt{\ell(\ell + d - 2) + \left(\frac{d-2}{2}\right)^2} = \ell + \frac{d-2}{2}.$$

In particular, for $d = 3$ the allowed energies are $\pm \omega_\ell = \pm(\ell + \frac{1}{2})$.

From now on, let us fix $d = 3$. The most general solution is a sum of all solutions of the form (22), namely

$$\chi(\sigma, \theta, \phi) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} (b_{\ell,m}^+ e^{\omega_\ell \sigma} + b_{\ell,m}^- e^{-\omega_\ell \sigma}) Y_{\ell,m}(\theta, \phi),$$

where $b_{\ell,m}^{\pm}$ are some operators acting on Hilbert space. The reality condition — using $Y_{\ell,m}^* = (-1)^m Y_{\ell,-m}$ — is

$$\begin{aligned}\chi(\sigma, \theta^i) &= \chi(-\sigma, \theta^i)^\dagger = \sum_{\ell,m} (-1)^m [(b_{\ell,m}^+)^\dagger e^{-\omega_\ell \sigma} + (b_{\ell,m}^-)^\dagger e^{\omega_\ell \sigma}] Y_{\ell,-m}(\theta, \phi) \\ &= \sum_{\ell,m} [(-1)^m (b_{\ell,-m}^-)^\dagger e^{\omega_\ell \sigma} + (-1)^m (b_{\ell,-m}^+)^\dagger e^{-\omega_\ell \sigma}] Y_{\ell,m}(\theta, \phi)\end{aligned}$$

(relabeling $m \rightarrow -m$ in passing from the first to the second line), so by imposing that this agrees with (25) for all σ and θ, ϕ , we learn that

$$b_{\ell,m}^+ = (-1)^m (b_{\ell,-m}^-)^\dagger.$$

It is more natural to redefine the modes as

$$a_{\ell,m}^+ = \sqrt{2\omega_\ell} b_{\ell,m}^+, \quad a_{\ell,m}^- = \sqrt{2\omega_\ell} (-1)^m b_{\ell,-m}^-$$

such that

$$a_{\ell,m}^+ = (a_{\ell,m}^-)^\dagger.$$

This also leads to a more standard expression for the field χ , namely

$$\chi(\sigma, \theta, \phi) = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \frac{1}{\sqrt{2\omega_\ell}} [a_{\ell,m}^+ e^{\omega_\ell \sigma} Y_{\ell,m}(\theta, \phi) + a_{\ell,m}^- e^{-\omega_\ell \sigma} Y_{\ell,m}(\theta, \phi)^*]$$

which is now manifestly hermitian.

(c) Computation of the Hamiltonian H

We are left to compute H , by integrating

$$H = \frac{1}{2} \int_{S^2} d\Omega \left[-(\partial_\sigma \chi)^2 + \chi \left(-\square_{S^2} + \frac{1}{4}(d-2)^2 \right) \chi \right].$$

For the first term, notice that

$$\partial_\sigma \chi = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \sqrt{\frac{\omega_\ell}{2}} [a_{\ell,m}^+ e^{\omega_\ell \sigma} Y_{\ell,m}(\theta, \phi) - a_{\ell,m}^- e^{-\omega_\ell \sigma} Y_{\ell,m}(\theta, \phi)^*]$$

such that

$$-\int_{S^2} (\partial_\sigma \chi)^2 = \sum_{\ell,m} \frac{\omega_\ell}{2} [-(-1)^m a_{\ell,m}^+ a_{\ell,-m}^+ e^{2\omega_\ell \sigma} - (-1)^m a_{\ell,m}^- a_{\ell,-m}^- e^{-2\omega_\ell \sigma} + a_{\ell,m}^+ a_{\ell,m}^- + a_{\ell,m}^- a_{\ell,m}^+].$$

To obtain this, we have used the orthogonality condition from the exercise. For the second part, we have instead

$$\left(-\square_{S^2} + \frac{1}{4}(d-2)^2 \right) \chi = \sum_{\ell=0}^{\infty} \sum_{m=-\ell}^{\ell} \frac{\omega_\ell^{3/2}}{\sqrt{2}} [a_{\ell,m}^+ e^{\omega_\ell \sigma} Y_{\ell,m}(\theta, \phi) + a_{\ell,m}^- e^{-\omega_\ell \sigma} Y_{\ell,m}(\theta, \phi)^*]$$

so

$$\int_{S^2} \chi \left(-\square_{S^2} + \frac{1}{4}(d-2)^2 \right) \chi = \sum_{\ell, m} \frac{\omega_\ell}{2} \left[(-1)^m a_{\ell, m}^+ a_{\ell, -m}^+ e^{2\omega_\ell \sigma} + (-1)^m a_{\ell, m}^- a_{\ell, -m}^- e^{-2\omega_\ell \sigma} + a_{\ell, m}^+ a_{\ell, m}^- + a_{\ell, m}^- a_{\ell, m}^+ \right]$$

Adding both contributions, we see that the $a^+ a^+$ and $a^- a^-$ terms cancel, but the $a^+ a^-$ and $a^- a^+$ terms survive, and

$$H = \sum_{\ell, m} \frac{\omega_\ell}{2} [a_{\ell, m}^+ a_{\ell, m}^- + a_{\ell, m}^- a_{\ell, m}^+].$$

In particular, we point to the fact that H does not depend on the time coordinate σ .

(d) **Computation of the Commutator** $[\chi, \partial_\sigma \chi]$

Since we have explicit expressions for χ and $\partial_\sigma \chi$, let us compute the commutator explicitly:

$$\begin{aligned} [\chi, \partial_\sigma \chi] &= \frac{1}{2} \sum_{\ell, m} \sum_{\ell', m'} \sqrt{\frac{\omega_{\ell'}}{\omega_\ell}} [a_{\ell, m}^+, a_{\ell', m'}^+] e^{(\omega_\ell + \omega_{\ell'})\sigma} Y_{\ell, m} Y_{\ell', m'} \\ &\quad - [a_{\ell, m}^+, a_{\ell', m'}^-] e^{(\omega_\ell - \omega_{\ell'})\sigma} Y_{\ell, m} Y_{\ell', m'}^* + [a_{\ell, m}^-, a_{\ell', m'}^+] e^{(-\omega_\ell + \omega_{\ell'})\sigma} Y_{\ell, m}^* Y_{\ell', m'} \\ &\quad - [a_{\ell, m}^-, a_{\ell', m'}^-] e^{-(\omega_\ell + \omega_{\ell'})\sigma} Y_{\ell, m}^* Y_{\ell', m'}^*. \end{aligned}$$

Since this must equal (14) from the exercise, we see in particular that the RHS cannot depend on σ . So we must have $[a_{\ell, m}^+, a_{\ell', m'}^+] = 0$, and by taking the hermitian conjugate of that also $[a_{\ell, m}^-, a_{\ell', m'}^-] = 0$. Looking at the second and third terms, we see that in addition we must have

$$[a_{\ell, m}^+, a_{\ell', m'}^-] \propto \delta_{\ell, \ell'}.$$

Using this, we rewrite the above equation as

$$[\chi, \partial_\sigma \chi] = \sum_{\ell, m, m'} [a_{\ell, m}^-, a_{\ell, m}^+] Y_{\ell, m}(\theta, \phi) Y_{\ell, m'}(\theta, \phi)^*.$$

After relabeling the dummy indices m, m' in one of the terms and using the property that $[A, B] = -[B, A]$, we get

$$[\chi, \partial_\sigma \chi] = \sum_{\ell, m} c(\ell, m) Y_{\ell, m}(\theta, \phi) Y_{\ell, m}(\theta, \phi)^*,$$

where $c(\ell, m)$ are some constants. Using the orthogonality relation for spherical harmonics,

$$\sum_{m=-\ell}^{\ell} Y_{\ell, m}(\theta, \phi) Y_{\ell, m}(\theta, \phi)^* = 1,$$

we find that

$$[\chi, \partial_\sigma \chi] = \sum_{\ell} c(\ell, m).$$

Matching this with the result from the exercise (equation 14), we conclude that $c(\ell, m) = 1$ for all ℓ, m , so

$$[a_{\ell, m}^-, a_{\ell', m'}^+] = \delta_{\ell, \ell'} \delta_{m, m'}.$$

(e) **Hamiltonian H in Normal-Ordered Form**

Returning to our expression for the Hamiltonian H , we get

$$H = \sum_{\ell,m} \frac{\omega_\ell}{2} (a_{\ell,m}^+ a_{\ell,m}^- + [a_{\ell,m}^-, a_{\ell,m}^+] + a_{\ell,m}^- a_{\ell,m}^+).$$

Since $[a_{\ell,m}^-, a_{\ell,m}^+] = 1$, we have

$$H = \sum_{\ell,m} \omega_\ell a_{\ell,m}^+ a_{\ell,m}^- + E_0,$$

where

$$E_0 = \sum_{\ell,m} \frac{\omega_\ell}{2} = \sum_{\ell=0}^{\infty} (\ell + \frac{1}{2})^2 = \infty.$$

The term E_0 is a divergent constant that we subtract from H . This subtraction is equivalent to normal-ordering H .

The Hamiltonian H becomes an infinite sum of harmonic oscillators, each associated with a creation operator $a_{\ell,m}^+$ and an annihilation operator $a_{\ell,m}^-$. To define our Hilbert space, we start with a vacuum state $|\Omega\rangle$ that obeys

$$\forall \ell, m : a_{\ell,m}^- |\Omega\rangle = 0.$$

We build excited states by acting with the creation operators $a_{\ell,m}^+$. For instance, the state $|\ell, m\rangle = a_{\ell,m}^+ |\Omega\rangle$ has energy

$$H|\ell, m\rangle = [H, a_{\ell,m}^+]|\Omega\rangle + a_{\ell,m}^+ H|\Omega\rangle = \omega_\ell |\ell, m\rangle.$$

A complete basis of states in the theory is given by

$$|\Psi\rangle = \prod_{\ell=0}^{\infty} \prod_{m=-\ell}^{\ell} (a_{\ell,m}^+)^{n_{\ell,m}} |\Omega\rangle,$$

where $n_{\ell,m} \in \mathbb{N}$. The state $|\Psi\rangle$ has energy

$$H|\Psi\rangle = E(\Psi)|\Psi\rangle, \quad E(\Psi) = n_{0,0}\omega_0 + \sum_{m=-1}^1 n_{1,m}\omega_1 + \dots + \sum_{m=-\ell}^{\ell} n_{\ell,m}\omega_\ell + \dots$$

or equivalently

$$E(\Psi) = \sum_{\ell=0}^{\infty} n_\ell (\ell + \frac{1}{2}), \quad n_\ell = \sum_{m=-\ell}^{\ell} n_{\ell,m}.$$

(f) **Matching with Flat-Space Theory**

In the flat-space theory, the free scalar $\phi(x)$ has dimension $\Delta_\phi = \frac{d-2}{2} = \frac{1}{2}$, setting $d \rightarrow 3$. Its descendants

$$\partial_{\mu_1} \cdots \partial_{\mu_n} \phi(x)$$

have dimension $\Delta = \Delta_\phi + n = n + \frac{1}{2}$. These scaling dimensions match the one-particle states

$$|\ell, m\rangle = a_{\ell, m}^+ |\Omega\rangle, \quad E = \ell + \frac{1}{2},$$

if we set $\ell = n$. The number of states with energy $n + \frac{1}{2}$ also agrees. For $n = 1$, there are three descendants $\partial_\mu \phi$, and at the same time, there are three states with energy $\frac{3}{2}$, namely $|1, -1\rangle, |1, 0\rangle, |1, 1\rangle$. At the next level $n = 2$, there are 5 different descendants: there are $\frac{1}{2}d(d+1) = 6$ operators of the form $\partial_\mu \partial_\nu \phi$, but

$$\square \phi = \sum_{\mu=1}^3 \partial_\mu^2 \phi = 0,$$

so the number of independent operators is only $6 - 1 = 5$. Similarly, there are $2\ell + 1 = 5$ states of the form $|2, m\rangle$, with $m = -2, \dots, 2$. This reasoning continues for higher levels. We can also consider operators with multiple fields, such as $O =: \phi^2 :$ with dimension $\Delta_O = 2\Delta_\phi = 1$. This operator corresponds to a two-particle state $(a_{0,0}^+)^2 |\Omega\rangle$ with energy $2\omega_0 = 1$. Using similar logic, we can match a general operator of the schematic form

$$O =: \partial^{j_1} \phi \partial^{j_2} \phi \dots \partial^{j_n} \phi :$$

with scaling dimension $\Delta_O = \frac{n}{2} + j_1 + j_2 + \dots + j_n$ to an n -particle state with energy $E = \Delta_O$.

(g) **Computation of the Integral $F(r)$**

We are instructed to compute the integral

$$F(r) = \int_{S^{d-1}} d\Omega r^{d-1} x^\mu r j_\mu(x) = - \int_{S^{d-1}} d\Omega r^{d-2} x^\mu x^\nu T_{\mu\nu}(x)$$

with

$$T_{\mu\nu} = \partial_\mu \phi \partial_\nu \phi - \frac{1}{2} g_{\mu\nu} (\partial_\rho \phi)^2 + \xi (g_{\mu\nu} \square - \partial_\mu \partial_\nu) \phi^2.$$

Let us work in flat space, with coordinates $\{\tau, n^\mu\}$ defined by $x^\mu = e^\tau n^\mu$. In particular,

$$x^\mu \partial_\mu f(x) = \partial_\tau f(\tau, n), \quad x^\mu x^\nu \partial_\mu \partial_\nu = \partial_\tau^2 - \partial_\tau.$$

Therefore,

$$x^\mu x^\nu T_{\mu\nu} = \frac{1}{2} (\partial_\tau \phi)^2 - \frac{1}{2} h^{ij} \partial_i \phi \partial_j \phi + \xi [e^{2\tau} \square - \partial_\tau^2 + \partial_\tau] \phi^2,$$

where we use that

$$ds^2 = e^{2\tau} [d\tau^2 + h_{ij} d\theta^i d\theta^j].$$

Now let us perform the field redefinition $\phi(\tau, n) = e^{-\frac{1}{2}(d-2)\tau} \chi(\tau, n)$ as before. Moreover, from (53) it follows that

$$e^{2\tau} \square = \square_{S^{d-1}} + \partial_\tau^2 + (d-2)\partial_\tau$$

when acting on scalar functions. The first term in $x^\mu x^\nu T_{\mu\nu}$ transforms as

$$e^{(d-2)\tau} \frac{1}{2} (\partial_\tau \phi)^2 = \frac{1}{2} (\partial_\tau \chi)^2 + \frac{1}{2} \left(\frac{d-2}{2} \right)^2 \chi^2 - \frac{d-2}{4} \partial_\tau (\chi^2),$$

whereas the last term gives

$$e^{(d-2)\tau} (e^{2\tau}\square - \partial_\tau^2 + \partial_\tau) \phi^2 = [-(d-1)(d-2) + (d-1)\partial_\tau + \square_{S^{d-1}}]\chi^2.$$

If $\xi = \xi_c$, the full stress tensor simplifies to

$$\xi = \xi_c : \quad -e^{(d-2)\tau} x^\mu x^\nu T_{\mu\nu} = -\frac{1}{2}(\partial_\tau\chi)^2 + \frac{1}{2}\chi(-\square_{S^{d-1}}\chi) + \frac{1}{2}\left(\frac{d-2}{2}\right)^2 \chi^2,$$

neglecting a boundary term $\nabla_i(\dots)$ that vanishes once integrated over the sphere. This expression matches the Hamiltonian density that we started with.

If $\xi \neq \xi_c$, there is an additional term:

$$-e^{(d-2)\tau} x^\mu x^\nu T_{\mu\nu} \supset (\xi - \xi_c)[(d-1)(d-2) - (d-1)\partial_\tau - \square_{S^{d-1}}]\chi^2.$$

This contributes to H , although the term vanishes after integration over S^{d-1} . The remaining terms contribute as a mass term and an interaction term $\chi\partial_\tau\chi \sim \chi\Pi$.